





# Modeling Sunlight in Gazebo for Vision-Based Applications Under Varying Light Conditions

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**Keywords:** Robotics, Vision, the Sun, Lighting Conditions, Simulation, Localization, Gazebo.

**Abstract:** Vision is one of the well-researched sensing abilities of robots. However, applying vision-based algorithms can be challenging when used in different environmental conditions. One such challenge in vision-based localization is dynamic lighting conditions. In this paper, we present a new Gazebo plugin that enables realistic illumination changes depending on a current Sun's position. A plugin's underlying algorithm takes into account various parameters, such as date, time, latitude, longitude, elevation, pressure, temperature, and atmospheric refraction. Virtual experiments demonstrated effectiveness of the proposed plugin, and the source code is available for free academic use.

## 1 INTRODUCTION

Vision is one of key sensing abilities and is actively researched in robotics. In the medical field, robots equipped with visual sensors can assist during surgical operations (Komatsu et al., 2024). In the industrial sector, vision-enabled robots can accelerate assembly processes (Duan et al., 2024). In transportation, robotic vision can be utilized for autonomous navigation (Artono et al., 2024). Autonomous navigation algorithms rely on localization algorithms, which, in turn, are influenced by operating conditions. One of these conditions is dynamic lighting, which changes throughout a day, presenting a challenge for localization algorithms to overcome (Piasco et al., 2021; Oishi et al., 2019).

Testing algorithms in all possible conditions can be challenging. In order to reduce software costs, algorithms could be verified in a simulation before being tested in a real-world setting. Yet, the simulation environment should be reproducible in the real world, i.e., it needs to be realistic enough.

Objectives of this paper are to simulate realistic

sunlight in the Gazebo simulator and to evaluate an impact of varying lighting conditions on vision-based localization algorithms. This work presents a newly developed solar position plugin for the Gazebo simulator and examines its effects on vision-based localization algorithms' performance.


## 2 RELATED WORK


### 2.1 Vision-Based Localization


Localization is not limited to vision-based algorithms. Some works use other algorithms for localization, e.g., 2D (Yue et al., 2024) or 3D structure-based approaches (Sarlin et al., 2021). This paper focuses on vision-based algorithms.


Typically, vision-based localization (Safin et al., 2020; Mingachev et al., 2020) employs visual information (such as multi-color or gray-color images) or a combination of visual information and supplementary data (such as depth or odometry information).

A variety of features can be used for vision-based localization. For example, in (Drupt et al., 2024) the simultaneous localization and mapping (SLAM) algorithm used ORB features and was designed for underwater robots. Similarly, in (Zhang et al., 2024; Ji

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et al., 2024), SLAM systems employed ORB features but aimed at ground robots operating in dynamic environments. In (Adkins et al., 2024), ORB features were used for ground robots engaged in long-term activities. Other works employed SIFT (Lowe, 2004), SURF (Bay et al., 2006), and BRISK (Leutenegger et al., 2011) features.

Additionally, challenges in localization include dynamic obstacles (e.g., cars, pedestrians, or animals), weather conditions (e.g., rain, snow, or fog), changing lighting conditions and others.

Dynamic obstacles can introduce additional uncertainty of a transformation estimated by a localization algorithm if not properly addressed. Weather conditions, such as rain, snow, and fog, can reduce visibility over large distances and may cause a sensor lens to become occluded with water droplets, limiting its field of view and/or distorting obtained images.

Lighting affects data acquired by visual sensors, often resulting in multiple different representations of a same location. Common solutions to this issue include a use of lighting-invariant image representations (Corke et al., 2013), estimating differences in sequential data (Milford and Wyeth, 2012), and mapping an environment over multiple sessions (Labbé and Michaud, 2022).

Effects of different lighting conditions on localization algorithms can vary from minor to major, depending on a disparity of the lighting conditions of a query image used for localization and a reference image used for mapping. For example, in (Irie et al., 2012), the FAB-MAP algorithm achieved a recall rate of 0.6% when applied to datasets with large lighting changes, compared to a 20.2% recall rate with datasets experiencing small lighting changes. Therefore, lighting conditions can be a crucial factor in achieving accurate localization.

Furthermore, vision-based detection and segmentation algorithms typically rely on visual features such as color, texture, and contrast, which can be drastically affected by illumination changes.

## 2.2 Light Simulation in Virtual Environments

Light simulation in virtual environments is commonly referred to as rendering. There are several rendering techniques, including ray casting (Roth, 1982), ray tracing (Shirley and Morley, 2008), radiosity (Goral et al., 1984; Nishita and Nakamae, 1984), path tracing (Lafortune, 1996), and others. These techniques generally use rays to simulate light propagation. In our research, experiments were conducted using the Cascaded Image Voxel Cone Tracing global illumination

approach, which is provided by the Gazebo simulator (Koenig and Howard, 2004).

The Gazebo simulator (Koenig and Howard, 2004) can be configured to use different graphics rendering engines, including OGRE, OGRE-next (Rojtberg, Pavel and Rogers, David and Streeting, Steve and others, 2024), or OptiX (Parker et al., 2010). The light is simulated according to techniques used in a selected graphics rendering engine. In this work, OGRE-next graphics rendering engine was used.

## 3 METHODOLOGY

The Gazebo simulator (Koenig and Howard, 2004) is an open-source software widely used in robotics and is compatible with the Robot Operating System (ROS). Gazebo is featured by a large community support and numerous ready-to-use plugins – extensions that allow modifying simulator’s behavior.

Additionally, Gazebo’s abstraction layer enables use of different rendering engines, making it a valuable tool for research of lightning conditions’ simulation.

### 3.1 Simulation of Light

To simulate realistic lighting in Gazebo, we implemented our own solar position plugin. This plugin creates a virtual *Sun* in the simulated environment, with its position calculated using the Solar Position Algorithm (SPA) (Reda and Andreas, 2004), which has an uncertainty of  $\pm 0.0003^\circ$ . The algorithm allows calculating solar zenith and azimuth angles within a time frame from the year 2000 B.C. to 6000 A.D. The algorithm consists of the steps that calculate: Julian and Julian Ephemeris Day, century, and millennium; Earth heliocentric longitude, latitude, and radius vector; geocentric longitude and latitude; nutation in longitude and obliquity; true obliquity of the ecliptic; aberration correction; apparent sun longitude and apparent sidereal time at Greenwich; geocentric and right ascension; geocentric sun declination; observer local hour angle; topocentric sun right ascension; topocentric local hour angle; topocentric zenith and azimuth angles.

The Solar Algorithm includes calculations for incidence angle and sunrise, transit, and sunset times. However, these were not used in our plugin, as they are not necessary for calculating a solar zenith angle, an azimuth angle, and a distance.

The solar position plugin uses the following parameters, as required by the SPA, to calculate a solar zenith angle, an azimuth angle, and a distance: a year,

a month, a day, an hour, a minute, a second, a time-zone, a latitude, a longitude, an elevation, an annual average local pressure, an annual average local temperature, an atmospheric refraction at a sunrise and a sunset, a fractional second difference between UTC and UT (Delta UT1), a difference between Earth's rotation time and terrestrial time (Delta T).

### 3.2 Vision-Based Localization Algorithms

In this paper, two approaches of localization algorithms were tested: a direct image matching and an image features matching. In the direct image matching, intensity values of images are directly compared. In the image features matching, sets of detected features are compared. In this study, SIFT, SURF, ORB, and BRISK features were tested. Feature extraction and description were performed using the OpenCV (version 4.5.4) library (Bradski, 2000). Using the SPA, three datasets were collected at different times of the day: at 4 AM, at 12 PM, at 8 PM.

### 3.3 Performance Metrics

To analyze effects of the solar position plugin, two metrics were used: a mean difference between datasets and a standard deviation between datasets. The mean difference between two datasets is calculated using the following equation:

$$E = \frac{1}{n} \sum_{i=1}^n D(A_i, B_i) \quad (1)$$

where  $E$  is the mean difference,  $n$  is a number of images in the dataset,  $D(A_i, B_i)$  is a difference between corresponding images, and  $A$  and  $B$  are the datasets.

The standard deviation between two datasets is calculated using the following equation:

$$s = \sqrt{\frac{1}{n-1} \sum_{i=1}^n (D(A_i, B_i) - E)^2} \quad (2)$$

where  $s$  is the standard deviation,  $n$  is the number of images in a dataset,  $D(A_i, B_i)$  is a difference between corresponding images in the datasets  $A$  and  $B$ ,  $A_i$  and  $B_i$  are the  $i$ -th images of the corresponding datasets, and  $E$  is the mean difference between the two datasets.

A difference between two images using the direct image matching approach is calculated using the following equation:

$$D = \frac{1}{wh} \sum_{x=1}^w \sum_{y=1}^h distance(a_{xy}, b_{xy}) \quad (3)$$

where  $D$  is the difference between the two images,  $w$  is a width of an image,  $h$  is a height of an image,  $distance(a_{xy}, b_{xy})$  is a distance function applied to pixel values, and  $a$  and  $b$  are images with  $a_{xy}$  and  $b_{xy}$  representing pixel values at corresponding coordinates.

A distance between two pixels using the boolean approach is calculated based on the following equation:

$$distance = \begin{cases} 0 & \text{if } a_{xy} == b_{xy} \\ 1 & \text{if } a_{xy} != b_{xy} \end{cases} \quad (4)$$

where  $a_{xy}$  and  $b_{xy}$  are pixel values.

A distance between two pixels in gray-scale images using the Euclidean approach is calculated based on the following equation:

$$distance = \frac{(a_{xy} - b_{xy})^2}{255^2} \quad (5)$$

where  $a_{xy}$  and  $b_{xy}$  are pixel values.

A distance between two pixels in RGB images using Euclidean approach is calculated based on the following equation:

$$distance = \frac{1}{255^2 * 3} \sum_{c=1}^3 (a_{xyc} - b_{xyc})^2 \quad (6)$$

where  $a_{xyc}$  and  $b_{xyc}$  are pixel values in channel  $c$ . RGB images contain three channels.

A difference between two images using the image features matching approach is calculated based on the following equation:

$$D = \frac{F_{total} - F_{matched}}{F_{total}} \quad (7)$$

where  $D$  represents a difference between the two images,  $F_{total}$  is a total number of features,  $F_{matched}$  is a number of features that were matched.

A total number of features is calculated using the following equation:

$$F_{total} = F_a + F_b - F_{matched} \quad (8)$$

where  $F_{total}$  is the total number of features,  $F_a$  is a number of features detected in image  $a$ ,  $F_b$  is a number of features detected in image  $b$ ,  $F_{matched}$  is a number of features that were matched.

Lowe's ratio was used for feature matching, meaning that features are matched if a distance of the closest match candidate is sufficiently different from that of the second closest match candidate. A threshold value was set to 0.5, meaning the closest match is accepted if it is at least twice as close as the next closest candidate.

### 3.4 Experimental Setup

To the best of our knowledge, there is no common standard environment for testing localization algorithms. While exploring existing environment models for our experiments in Gazebo, we noted that some models lack a roof (which allows a sunlight to leak from above), other models are too simple (containing only a few objects) or too complex to be processed quickly by a typical PC. Finally, a model (Nicolas-3D, 2024) shown in Fig. 1 was selected.

Lighting in the Gazebo simulator can be improved using global illumination techniques, which typically yield more realistic results. Currently, the Gazebo simulator supports only Voxel Cone Tracing (VCT) and Cascaded Image Voxel Cone Tracing (CIVCT). Both approaches were tested in the proposed environment, and the CIVCT approach was chosen because it subjectively produced more visually realistic results.

The source code of the Gazebo simulator was modified to use CIVCT for global illumination. These modifications were necessary because the CIVCT approach requires binding a camera, and at the time, no method was found to bind a non-GUI camera other than modifying the source code directly. The modified source code is available on GitLab<sup>1</sup>.

The chosen trajectory of the camera used in the experiments is marked in red in Fig. 2. Camera poses were calculated along this trajectory at 1-meter intervals, except cases where the camera performs a turn; in those instances, the pose was set at the beginning of the next line segment. The camera was always pointed toward the next trajectory point. As a result, each dataset contained 991 images, which were manually checked for consistency.

A PC with Intel Core i3 processor, 8 GB of RAM, and Nvidia GeForce GTX 950 GPU was used for the evaluation. The simulation included the testing environment, the camera with a resolution of 640x480 pixels, and the proposed solar position plugin. The parameters used during the experiments are presented in Table 1.

## 4 RESULTS

Tables 2 and 3 present results of the experiments. Table 2 shows mean difference values for each algorithm, while Table 3 provides standard deviation values. In both tables, "4-12," "4-20," and "12-20" correspond to matching pairs of images acquired at 4:00

<sup>1</sup><https://gitlab.com/lirs-kfu/gazebo-solar-position-plugin/-/tree/main/gz-sim>

Table 1: Parameters of the solar position plugin used in the experiments.

Parameter	Value
Year	2024
Month	7
Day	1
Hour	4, 12, or 20
Minute	0
Second	0
Timezone	+3 hours
Latitude	+55.792115°
Longitude	+49.122254°
Elevation	90 meters
Pressure	1013.25 millibars
Temperature	0° Celsius
Atmospheric refraction	0.5667
Delta UT1	0 seconds
Delta T	69.184 seconds

AM vs 12:00 PM, 4:00 AM vs 8:00 PM, and 12:00 PM vs 8:00 PM, respectively.

Table 2: Mean difference values in the experiments.

Algorithm	4-12	4-20	12-20
Boolean-rgb	0.70159	0.69474	0.70102
Boolean-grey	0.70078	0.66502	0.70063
Euclidean-rgb	0.06425	0.00308	0.07117
Euclidean-grey	0.06308	0.00295	0.07000
SIFT	0.84577	0.63501	0.83832
SURF	0.94056	0.81301	0.93860
ORB	0.97809	0.89677	0.97376
BRISK	0.95350	0.86339	0.95288

Table 3: Standard deviation of differences in the experiments.

Algorithm	4-12	4-20	12-20
Boolean-rgb	0.06709	0.07227	0.06767
Boolean-grey	0.06680	0.10038	0.06752
Euclidean-rgb	0.00917	0.00190	0.01036
Euclidean-grey	0.00862	0.00176	0.01014
SIFT	0.04907	0.14882	0.05197
SURF	0.01816	0.10936	0.02048
ORB	0.01723	0.09651	0.02059
BRISK	0.01580	0.09893	0.01765

Tables 4 and 5 present the results of the tests. Table 4 shows the mean difference values for each algorithm, while Table 5 provides the standard deviation values. In both tables, "4-4," "12-12," and "20-20" correspond to tests using images acquired at 4:00 AM, 12:00 PM, and 8:00 PM, respectively, to evaluate the deterministic behavior of the algorithms.

Equation 4 was used to calculate a difference between images in the *boolean-rgb* algorithm and in the



Figure 1: Virtual environment images. The left column shows morning (4 AM) images, the middle column shows day (12 PM) images, and the right column shows evening (8 PM) images. Rows correspond to images with IDs 264, 394, 573, 741, and 800.

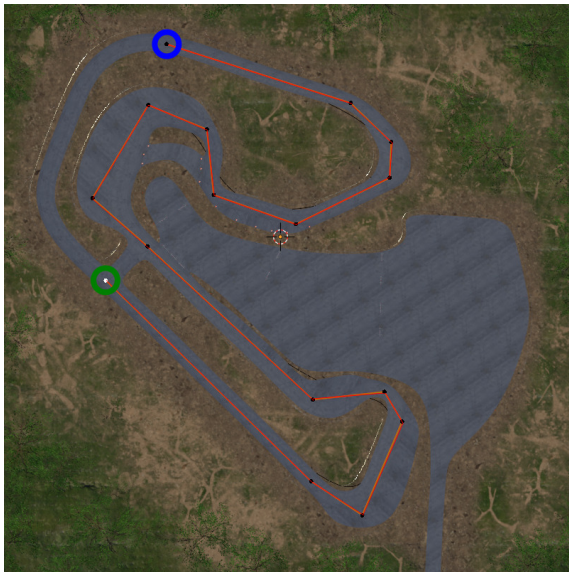


Figure 2: Trajectory of the camera used in the experiments is marked in red. The camera starts at the point marked with the blue circle and finishes at the point marked with the green circle.

Table 4: Mean difference values in the tests.

Algorithm	4-4	12-12	20-20
Boolean-rgb	0	0	0
Boolean-grey	0	0	0
Euclidean-rgb	0	0	0
Euclidean-grey	0	0	0
SIFT	1.19e-5	0	4.54e-6
SURF	0	0	0
ORB	8.04e-6	0	8.04e-6
BRISK	0.01224	0.00751	0.01274

*boolean-grey* algorithm. Equation 6 was used to calculate a difference between images in the *Euclidean-rgb* algorithm. Equation 5 was used to calculate the difference between images in the *Euclidean-grey* algorithm. For SIFT, SURF, ORB, and BRISK, the default parameters provided by the OpenCV library were used.

## 5 DISCUSSION

### 5.1 Limitations

In the experiments, the Sun's position was not adjusted according to the simulation time. However, with some source code modifications, the solar position plugin can be configured to move the Sun based on simulation time. These modifications are expected to be included in future versions of the plugin.

Table 5: Standard deviation of differences in the tests.

Algorithm	4-4	12-12	20-20
Boolean-rgb	0	0	0
Boolean-grey	0	0	0
Euclidean-rgb	0	0	0
Euclidean-grey	0	0	0
SIFT	0.00023	0	0.00014
SURF	0	0	0
ORB	0.00025	0	0.00025
BRISK	0.00655	0.00324	0.00578

The background color (sky color) in the experiments was set to grey. Unfortunately, the sky provided by the Gazebo simulator was not influenced by the angle of sunlight; for instance, it remained mostly blue when it should have been red. As a result, the default sky was not used in the experiments. This limitation should be considered when interpreting the experimental results, as the sky color remained constant and did not affect any of the algorithms. The sky occupied up to 50% of the image area (as the camera was parallel to the ground), which suggests that the algorithms might be more significantly affected if the sky was adjusted according to the angle of sunlight. To achieve more realistic results, a sky model that responds to the sunlight angle is necessary. Additionally, the Gazebo simulator lacks a night sky model. Incorporating a night sky, with the Moon and stars, could improve accuracy of results during nighttime and possibly during twilight.

The experiments did not include the robot's shadows. Shadows were turned off because only a camera was used, and the shadows cast by the camera alone did not appear realistic. The shadow area of a robot in an image can range from small to large. Since robot shadows are dependent on the angle of sunlight, enabling them in future experiments could result in significant image differences at different times of the day.

It was observed during the simulation that sunlight could pass through the virtual ground plane from below and illuminate objects above it. The model used in the experiments did not include any textures beneath the ground, which may have resulted in slightly less realistic outcomes. It is believed that using light-absorbing textures on the underside of the virtual ground plane could help eliminating some unrealistic effects, such as objects being lit from below by a Sun.

Default parameters provided by the OpenCV were used for feature detection and description. As a result, it is unclear whether the optimal parameters were used and how these may have affected the results.

## 5.2 Analysis

Despite all the factors mentioned above, the experimental results clearly demonstrated the effects of sunlight on the algorithms. Notations from Table 6 are used in analysis presented in this subsection.

Table 6: Notations for results' analysis.

Value	Notation
Mean difference between morning images and day images	$E_{md}$
Mean difference between morning images and evening images	$E_{me}$
Mean difference between day images and evening images	$E_{de}$
Standard deviation of difference between morning images and day images	$s_{md}$
Standard deviation of difference between morning images and evening images	$s_{me}$
Standard deviation of difference between day images and evening images	$s_{de}$

It can be observed that  $E_{me}$  is less than  $E_{md}$  or  $E_{de}$ . The morning images were taken approximately one hour after the sunrise, while the evening images were taken about half an hour before the sunset. As a result, the morning and evening images were darker compared to those taken in the middle of the day.

The Euclidean algorithms demonstrated differences between images using the Euclidean distance. It can be observed that  $E_{de}$  is slightly larger than  $E_{md}$ . This can be explained by a fact that the evening images were taken only half an hour before the sunset, while the morning images were taken an hour after the sunrise. Therefore, the morning conditions were closer to the day conditions. However, this was not the case for the boolean algorithms, where  $E_{md}$  and  $E_{de}$  are approximately the same. It is believed that the boolean algorithms do not handle image matching well enough and, as a result, produce poorer results.

The difference measured using feature-based algorithms is not expected to be fully accurate because these algorithms do not account for ground-truth relationships between features in images from different datasets. Unfortunately, manual dataset labeling was not possible at the time. Nevertheless, the results of feature-based algorithms still show correlations with each other. It can be observed that the SIFT algorithm performed well, while the SURF, BRISK, and ORB algorithms performed worse. The poor results may be attributed to a lack of experience with these fea-

tures, as only the default parameters provided by the OpenCV library were used, and no parameter analysis was conducted.

As for the standard deviation results,  $s_{me}$  is the largest, except for the Euclidean algorithms. The large value of  $s_{me}$  indicates that, even though the mean difference between images was not significant, some individual images exhibited more variation than others. This suggests that the morning and evening images may not be interchangeable. Although darkness levels in the morning and evening images are approximately the same, the angle of sunlight and shadows still differ, leading to the greater deviation. For the Euclidean algorithms,  $s_{me}$  is smaller than  $s_{md}$  and  $s_{de}$ , indicating that the darkness levels in the morning and evening images are similar, while the day images are brighter.

The experimental results indicated that SIFT, ORB, and BRISK produce different features for the same dataset across different runs. Since the SURF algorithm does not produce noticeable differences for a single dataset, and the source code of the experiments was thoroughly reviewed, it was concluded that SIFT, ORB, and BRISK are either not deterministic algorithms or are not properly configured to be deterministic. As a result, the differences between datasets may vary slightly.

## 6 CONCLUSIONS

In this paper, the solar position plugin for the Gazebo simulator was presented. The Sun's position in the sky depends not only on gravitational forces but also on light propagation parameters, such as a speed of light in a particular environment, atmospheric conditions, and time-related errors. Therefore, the SPA was used to accurately model the Sun's position in the virtual environment. Virtual experiments with the developed solar position plugin demonstrated that SIFT, SURF, ORB, and BRISK algorithms were affected by dynamic lighting. The plugin is available for free academic use at Gitlab account of our Laboratory of Intelligent Robotic Systems (LIRS)<sup>2</sup>.

## ACKNOWLEDGEMENTS

This paper has been supported by the Kazan Federal University Strategic Academic Leadership Program ("PRIORITY-2030").

<sup>2</sup>Laboratory of Intelligent Robotic Systems, GitLab, <https://gitlab.com/lirs-kfu/gazebo-solar-position-plugin>

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